

Lawrence Berkeley National Laboratory

Lawrence Berkeley National Laboratory

Title

An alkaline spring system within the Del Puerto ophiolite (California USA): A Mars analog site

Permalink

<https://escholarship.org/uc/item/30d990x8>

Author

Blank, J.G.

Publication Date

2009-03-02

1 **An Alkaline Spring System within the Del Puerto Ophiolite (California USA): A**
2 **Mars Analog Site**

3
4

5 J.G. Blank (1, 2*), S. Green (2†), D. Blake (2), J.W. Valley (3), N.T. Kita (3), A. Treiman
6 (4), P.F. Dobson (5)

7

8 (1) SETI Institute, Mountain View CA 94043 USA; (2) NASA/Ames Research Center,
9 Moffett Field CA 94035 USA; (3) Department of Geology and Geophysics, University of
10 Wisconsin, Madison, WI 53706 USA; (4) Lunar and Planetary Institute, Houston, TX
11 77058 USA; (5) Earth Sciences Division, Lawrence Berkeley National Laboratory,
12 Berkeley CA 94720 USA

13

14 Key words: Mars analog, dolomite, alkaline springs, biosignature

15

16 **Abstract**

17 Mars appears to have experienced little compositional differentiation of primitive
18 lithosphere, and thus much of the surface of Mars is covered by mafic lavas. On Earth,
19 mafic and ultramafic rocks present in ophiolites, oceanic crust and upper mantle that have
20 been obducted onto land, are therefore good analogs for Mars. The characteristic
21 mineralogy, aqueous geochemistry, and microbial communities of cold-water alkaline
22 springs associated with these mafic and ultramafic rocks represent a particularly

* Corresponding author. Tel. 1 650 810 0232. E-mail address: jblank@seti.org

† Now at Department of Oceanography, Florida State University, Tallahassee, FL 32306 USA

23 compelling analog for potential life-bearing systems. Serpentinization, the reaction of
24 water with mafic minerals such as olivine and pyroxene, yields fluids with unusual
25 chemistry (Mg-OH and Ca-OH waters with pH values up to ~12), as well as heat and
26 hydrogen gas that can sustain subsurface, chemosynthetic ecosystems. The recent
27 observation of seeps from pole-facing crater and canyon walls in the higher Martian
28 latitudes supports the hypothesis that even present conditions might allow for a rock-
29 hosted chemosynthetic biosphere in near-surface regions of the Martian crust. The
30 generation of methane within a zone of active serpentinization, through either abiogenic
31 or biogenic processes, could account for the presence of methane detected in the Martian
32 atmosphere. For all of these reasons, studies of terrestrial alkaline springs associated
33 with mafic and ultramafic rocks are particularly timely. This study focuses on the
34 alkaline Adobe Springs, emanating from mafic and ultramafic rocks of the California
35 Coast Range, where a community of novel bacteria is associated with the precipitation of
36 Mg-Ca carbonate cements. The carbonates may serve as a biosignature that could be
37 used in the search for evidence of life on Mars.

38

39 **1. Introduction**

40

41 A critical challenge facing the search for life in the solar system is the identification of
42 unambiguous evidence of life (cf., Beaty et al., 2005). The presence of microbial life on
43 Earth or an extraterrestrial planet does not ensure our ability to detect it. Evidence of life
44 must be distinctive from a landscape created by abiotic processes (cf., Dietrich and
45 Perron, 2006). The presence of water is deemed to be one of the key requirements for

46 identifying an environment capable of hosting life on Mars (e.g., Knoll and Grotzinger,
47 2006). The goal of this study is to identify possible biosignatures from a Martian analog
48 environment, namely, alkaline springs associated with ophiolites, sections of ocean crust
49 and upper mantle that have been obducted onto continental crust, experiencing varying
50 degrees of hydrothermal alteration in the process.

51

52 Serpentinization, the reaction of water with olivine- and pyroxene-rich rocks common in
53 mafic and ultramafic rocks to form serpentine, also produces heat and hydrogen gas that
54 can sustain subsurface, chemosynthetic ecosystems, and also results in the formation of
55 Mg-rich alkaline fluids. These fluids, when mixed with seawater (as seen at Lost City;
56 Kelley et al., 2005) or emanating as surface waters (e.g., as described comprehensively
57 by Pentecost, 2005) can produce substantial volumes of secondary carbonate deposits
58 (e.g., Surour and Arafa, 1997). Alkaline springs associated with mafic and ultramafic
59 rocks are model settings in which to identify possible mechanisms of biosignature
60 formation because these compositions of rocks have persisted throughout all of the
61 Earth's history. More importantly, low-temperature aqueous alteration processes (such as
62 serpentinization) associated with mafic and ultramafic rocks on Earth are thought to be
63 geologically similar to those occurring on Mars (e.g., Boston et al., 1992; Ming et al.,
64 2006; Wyatt and McSween, 2006).

65

66 *1.1 Mafic and ultramafic rocks as analog settings for early Earth, early Mars, and*
67 *other rocky planets*

68

69 Interaction between reducing rocks (e.g., unweathered basalts and ultramafic rocks) and
70 water results in an exothermic reaction that also produces hydrogen and methane, both
71 potential energy sources for chemosynthetic microorganisms (Kelley et al., 2005; Sleep
72 et al., 2004). Unfortunately, more detailed characterization of these systems is often
73 limited by their relative inaccessibility - whether in the deep-sea hydrothermal
74 environments or deep within the continental crust. More accessible systems are offered
75 by ophiolite terranes, sections of oceanic crust and upper mantle that have been obducted
76 onto land and which include both basaltic and ultramafic rocks. Similar rock types are
77 (and were) abundant on planetary bodies - the crusts of differentiated bodies (such as
78 Earth, Mars, Venus, and 4 Vesta) contain basaltic and ultramafic rock, and most
79 undifferentiated bodies (chondritic asteroids) are composed entirely of ultramafic rocks.

80

81 The serpentinization of mafic and ultramafic rocks can provide reduced substrates
82 suitable for microbial growth, and can yield secondary phases that may act as a
83 preservation medium for microbial organisms and their biosignatures (Fisk and
84 Giovannoni, 1999). The liberation of H₂ in these systems by mineral-water interaction
85 may be partially self-sustaining, given that a volume increase of as much as 60% during
86 serpentinization (Shervais et al., 2005a) creates the potential for mechanical fracturing,
87 which continually exposes new, unreacted mineral surfaces to water and, potentially,
88 organisms. Such an environment can persist for long periods, as the heat generated by
89 serpentinization has been shown to be sufficient to drive hydrothermal circulation of
90 highly reducing fluids over tens of thousands of years (Früh-Green et al., 2003).

91

92 On the early Earth, mafic and ultramafic rocks occurring in oceanic-type crust were
93 abundant, but little of this ancient crust remains today in a form that has not been highly
94 altered. Where present, obducted mafic and ultramafic rocks associated with ophiolite
95 terranes may represent an excellent terrestrial analog to Martian geology, since identified
96 Martian meteorites are either basalts or ultramafic rocks (e.g., Singer and McSween,
97 1993), and recent mapping of the Mars surface has revealed the dominance of mafic
98 rocks (Christensen et al., 2005). Any aqueous alteration of the Martian surface would
99 thus involve interaction with mafic and ultramafic rocks. This hypothesis is supported by
100 evidence from Martian meteorites, in which the predominant style of aqueous alteration
101 is that of olivine to phyllosilicates (Newsom et al., 2001; Treiman and Goodrich, 2002;
102 Leshin and Vicenzi, 2006), and to carbonates (Treiman and Romanek, 1998; Leshin et
103 al., 1998; Eiler et al., 2002), analogous to serpentinization of ophiolites. Additionally, the
104 recent discovery of hematite at the Meridiani Planum on Mars (Squyres et al., 2004a, b)
105 and quartz veinlets in eucrite meteorites (Treiman et al., 2004) and in a Mars meteorite
106 (Valley et al., 1997) are indicative of a history of aqueous alteration and activity on the
107 surface of Mars and other planetary bodies (e.g., asteroids). The generation of methane
108 within a zone of active serpentinization on Mars (through either abiogenic or biogenic
109 processes) could account for the presence of methane detected in the Martian atmosphere
110 (Formisano et al., 2004). Although currently Earth is the only planet we know of where
111 liquid water is stable at the surface, models based on recent satellite and Mars rover
112 observations of aeolian and fluvial sediments (e.g., Baker, 2006; Andrews-Hanna et al.,
113 2007) conclude that water was once present at the Martian surface, implying that both
114 surface and subsurface environments could have undergone serpentinization reactions,

115 and potentially supported life. While carbonates have not been identified on the surface
116 of Mars to date (although their presence is suggested by early returns from the Phoenix
117 Mars Lander), and recent detection of jarosite and other sulfate minerals hints that
118 portions of the surface of Mars are acidic today (Squyres and Knoll, 2005; Squyres et al.,
119 2006), carbonates may have been present at the surface of Mars early in this planet's
120 history (e.g., Treiman, 1998; Treiman and Romanek, 1998; Eiler et al., 2002), when more
121 widespread fluvial activity occurred (e.g., McEwan et al., 2007).

122

123 The continental borderland of California contains numerous ophiolite blocks of similar
124 age, ranging from ~172–164 Ma (Shervais et al., 2005b). Groundwaters circulating
125 within a number of ophiolite bodies found in the California Coast Range have reacted,
126 and continue to react, with the ultramafic rocks to yield cold springs with unusual
127 chemistry (Mg-OH and Ca-OH waters with pH values up to ~12; e.g., Barnes and O'Neil,
128 1971). Schulte et al. (2006) describe the petrology and mineral chemistry of the
129 ophiolite-hosted Complexion Spring (pH ~ 12), and have proposed criteria for identifying
130 serpentinized mafic rocks on Mars that may sustain chemosynthetic life. While such
131 waters can support a significant microbial load (Sleep et al., 2004), the springs and their
132 associated carbonate cements have not been studied in the context of biosignature
133 formation. The characteristic mineralogy and aqueous geochemistry of ophiolite-hosted
134 alkaline springs suggest that they may represent a particularly compelling analog for
135 potential life-bearing systems on early or modern Mars, and on the early Earth. For all of
136 these reasons, studies of terrestrial ophiolite-hosted alkaline springs and their associated
137 biota and secondary minerals are particularly timely.

138

139 *1.2 Del Puerto Ophiolite, California Coast Range*

140

141 Our field area is located within the Del Puerto Ophiolite, approximately 100 km SE of
142 San Francisco. The ophiolite is part of the California Coast Range and is Jurassic in age
143 (Evarts et al., 1992; Shervais et al., 2005b). The area is marked by rugged, sparsely
144 vegetated terrain, and outcrops exhibit extensive hydrothermal alteration (Evarts and
145 Schiffman, 1983). The ophiolite has been mapped as three distinct rock units: a basal
146 alpine peridotite member, a middle plutonic member, and an upper volcanic member
147 (Evarts, 1977). The study area and surrounding drainage system is hosted within the
148 peridotite body. Del Puerto Creek, the principal drainage for this region, flows eastward
149 toward the San Joaquin Valley. Adobe Springs are low-flow-rate features that discharge
150 into Adobe Creek, a tributary of Del Puerto Creek. The water in the creeks is a mixture
151 of seasonal surface run-off and local spring water.

152

153 Previously, two distinct alkaline water compositions were identified at the Adobe Spring
154 site: a high-pH (~12) Ca-OH water interpreted by Barnes et al. (1967) as evidence of
155 active serpentinization, and an alkaline (pH ~9) Mg-OH water interpreted to be a mixture
156 of ultramafic-derived and meteoric waters (Barnes and O'Neil, 1971). Barnes and O'Neil
157 (1971), O'Neil and Barnes (1971), and Blake and Peacor (1985) noted the presence of
158 calcite and dolomite cements in the drainages where these alkaline waters occur.

159

160 The high pH Ca-OH springs reported by Barnes et al. (1967) are no longer active at the
161 Adobe Springs site. However, the Mg-OH waters, which emanate from Adobe Springs
162 and are also present in the Del Puerto Creek and Adobe Creek drainages, appear to be the
163 source of the carbonate cements that line the drainages. A well drilled into the hillside
164 adjacent to Adobe Springs also produces moderately alkaline Mg-OH water, which is
165 bottled and sold for its reputed medicinal benefits (www.mgwater.com).

166

167 Initial research at this site has focused on characterizing and understanding the micron-
168 scale mineral, morphological and/or stable isotopic biosignatures in carbonate cements
169 associated with ophiolite-hosted alkaline springs. Detection of diagnostic biosignatures
170 would serve to suggest technologies or methodologies most useful for identifying past or
171 presently habitable zones on Mars during flight or sample-return missions. In addition,
172 characterization of the link between precipitating carbonate cements and microbial
173 activity within an ophiolitic terrain increases our understanding of the phylogeny and
174 physiology of microorganisms, including extremophiles, whose characteristics may
175 reflect the nature of primitive environments.

176

177 **2. Methods**

178

179 Water, rock, and microbial samples were collected in 2006 and 2007 from the drainage
180 area within a few hundred meters of Adobe Springs, near the confluence of Del Puerto
181 and Adobe Creeks (Figure 1). Water samples were periodically collected at three sample
182 sites near the confluence of the Del Puerto and Adobe Creeks: (1) Del Puerto Creek (the

183 main drainage within the Del Puerto Ophiolite), (2) Adobe Creek (a tributary of Del
184 Puerto Creek), which has intermittent flow, and (3) Adobe Springs well water (Figure 1).
185 Field measurements of pH and water temperature were recorded. Water samples were
186 filtered using a 0.45 micron filter and were kept cold prior to analysis. Water chemistry
187 analyses were performed by BC Laboratories (Bakersfield CA). Oxygen and hydrogen
188 isotopic analyses were conducted by the UC Berkeley Laboratory for Environmental and
189 Sedimentary Geochemistry. SOLVEQ (Reed, 1982), a computer program developed to
190 compute aqueous-mineral-gas equilibria, was used to determine mineral saturation
191 indices using measured Mg-OH water compositions.

192

193 Carbonate samples were collected at the two creek sites (Figures 1 and 2) for
194 petrographic and chemical analysis. Petrographic characterization of the cements was
195 conducted at Lawrence Berkeley National Laboratory. Selected carbonate samples were
196 analyzed for major and trace elements using an electron microprobe (EMP) at
197 NASA/Johnston Space Center. SEM images were collected at NASA/Ames and
198 NASA/JSC. Analysis of O-isotope variations in the cements on a microscopic scale was
199 conducted using a Secondary Ion Mass Spectrometry (SIMS) CAMECA ims-1280
200 instrument at the University of Wisconsin (Kita et al., 2007; Page et al., 2007; Blank et
201 al., 2007; Bowman et al., 2008). Instrumental bias of SIMS analysis is corrected by the
202 measurements of calcite and dolomite isotope standards and the spot to spot precision of
203 these in situ analyses is typically 0.3‰. (2σ).

204

205 A variety of water samples (well water, Adobe and Del Puerto Creek waters) and
206 microbial mat and sediment samples were collected for biologic characterization.

207 Genomic DNA was extracted from water samples using a commercial DNA extraction
208 procedure (Mo Bio Laboratories, Carlsbad, CA) after an initial filtration of the water
209 through a 0.2 micron filter. DNA was extracted from microbial mat and sediment
210 samples using a modified bead-beating method developed and tested in our laboratory
211 (Green et al., 2008). Samples were PCR-amplified with a variety of primer sets targeted
212 to ribosomal RNA (rRNA) genes of bacteria and Archaea, as well as functional genes for
213 sulfate-reducing prokaryotes (dissimilatory sulfite reductase, *dsrAB*) (Muyzer et al.,
214 1993; Muzyer and Smalla, 1998; Casamayor et al., 2002; Geets et al., 2006) and
215 methanogens (Methyl Coenzyme M Reductase A, *mcrA*) (Luton et al. 2002). Bacterial
216 and cyanobacterial primer sets (Muzyer et al., 1993; Muzyer and Smalla, 1998; Nubel et
217 al., 1997) were utilized for rapid community structure analysis using denaturing gradient
218 gel electrophoresis (DGGE).

219

220 **3. Results**

221

222 *Water chemistry*

223

224 As noted earlier, the highly alkaline (pH ~12) Ca-OH springs described by Barnes et al.
225 (1967) are no longer active, so sampling was confined to the Mg-OH alkaline waters
226 found in the well, springs, and creeks near Adobe Springs. During the dry summer
227 months, the only flows in this region are those fed by springs, and surface flow is
228 intermittent. Geochemical results of analyses of water samples collected from the Adobe
229 Creek well and Del Puerto Creek are presented in Table 1. Calculated log (Q/K) values
230 for disordered dolomite (1.76 and 2.30) and calcite (0.27 and 0.61) for the Adobe Creek

231 well water and Del Puerto Creek water, respectively, are positive, indicating that the Mg-
232 OH waters are supersaturated with respect to these carbonate phases. However, previous
233 studies have noted that precipitation of dolomite under ambient conditions is inhibited by
234 kinetic factors (e.g., Land, 1998).

235

236 *Carbonate Cements*

237

238 Carbonate cements line the creek beds, producing a conglomerate with clasts of
239 carbonate and fragments of eroded peridotite that range from sub-millimeter to tens-of-
240 cm in size. Initial investigations of the carbonate cements (Blank et al., 2006) have
241 revealed at least three distinct cement textures: laminated cements, massive or hummocky
242 cements, and dentate calcite crystals lining open pore space. Electron microprobe
243 analysis (Figure 3) indicates that the carbonates range in composition from dolomite to
244 calcite.

245

246 We detected $\delta^{18}\text{O}$ compositions for laminated carbonate ranging from 19.8–25.4 ‰_{VSMOW}
247 over a ~500 μm transect perpendicular to a serpentinite fragment grain boundary (Figure
248 4). For these samples, $\delta^{18}\text{O}$ values generally increase with increasing Mg content in the
249 carbonates, consistent with the observation by Tarutani et al (1969) that magnesian
250 calcites have a larger isotope fractionation relative to water than pure calcite. The range
251 in isotopic composition is consistent with compositions ($\delta^{18}\text{O} = 23.9\text{--}25.2\text{ ‰}$) for 3 bulk
252 Ca-Mg carbonate samples from Del Puerto Creek reported by Barnes and O'Neil (1971).
253 These bulk samples also exhibited a similar positive correlation between Mg content and

254 oxygen isotopic composition. We also observed variations in $\delta^{18}\text{O}$ values along-strike
255 within individual bands, with a variation of 1.2 ‰ encountered within a single ~50 μm -
256 thick dolomite band. This within band variability is significantly larger than analytical
257 uncertainty and attests to the heterogeneous environment of carbonate deposition. Using
258 the measured oxygen isotopic compositions and temperatures of the sampled waters and
259 dolomite-water and calcite-water oxygen isotope fractionation curves for both abiotic and
260 biotic systems (Tarutani et al., 1969; Schmidt et al., 2005; Vasconcelos et al., 2005;
261 O'Neil et al., 1969; Horita and Clayton, 2007), dolomite $\delta^{18}\text{O}$ values ranging from 24.3
262 to 26.6‰ and calcite $\delta^{18}\text{O}$ values ranging from 21.3 to 22.3‰ were calculated (Table 2).
263 The dolomite $\delta^{18}\text{O}$ values determined using the microbially mediated fractionation factor
264 of Vasconcelos et al. (2005) are 0.4 to 1.8‰ lower than those calculated using the
265 abiogenic fractionation factors of Tarutani et al. (1969) and Schmidt et al. (2005). There
266 is close agreement between the calcite isotopic compositions calculated using O'Neil et
267 al. (1969) (as modified in Friedman and O'Neil, 1977) and Horita and Clayton (2007). In
268 general, the dolomite $\delta^{18}\text{O}$ values calculated using the Vasconcelos fractionation equation
269 more closely match the measured $\delta^{18}\text{O}$ values obtained for the dolomitic portions of the
270 laminated carbonates, suggesting that dolomite precipitation at Adobe Springs was
271 microbially mediated.

272

273 *Microbial Communities*

274

275 A small-scale analysis of the microbiology of ophiolite-hosted waters was conducted
276 during the summers of 2006 and 2007. All water samples had a pH of approximately 9, as

277 did the water overlying the microbial mats. At the Adobe Springs field site, there are a
278 variety of different photosynthetically driven microbial communities, ranging from
279 laminated microbial mats to amorphous algal conglomerates, or periphyton (Figure 2).
280 Because of the ephemeral, and presumably seasonal, presence of these photosynthetic
281 communities, we have not yet ascertained their relationship to the deposition of the Ca-
282 Mg carbonate cements. However, the presence of laminated microbial mats in this
283 alkaline environment is a peculiar phenomenon that merits further investigation.

284

285 A clone library of approximately 150 16S rRNA gene sequences was generated from
286 three distinct microbial mat layers as well as water overlying the mat, from nearby water
287 wells, and from Del Puerto Creek water. The microbial mat clone library, composed of
288 75 sequences, reveals a diverse microbial community dominated by Cyanobacteria
289 (40%), Proteobacteria (27%), Bacteroidetes (13%) and Firmicutes (11%). Most of the
290 cyanobacterial sequences belong to two novel lineages of cyanobacteria, a finding
291 confirmed by the recovery of near full length rRNA gene sequences (Genbank accession
292 numbers [EU255702-EU255722](#); www.ncbi.nih.gov). These cyanobacterial sequences
293 belong to the order Oscillatoriales (filamentous, nonheterocystous cyanobacteria) and are
294 most similar to cyanobacterial sequences detected other in freshwater or brackish
295 microbial mats.

296

297 The most abundant bacterial phylum detected in the clone library generated from the
298 water samples is the phylum Bacteroidetes. In the Del Puerto Creek water, the microbial
299 community is dominated by a single species of Bacteroidetes (13 sequences of 29 total)

300 most closely related to the organism *Chimaericella alkaliphila*, a species isolated from a
301 highly alkaline (pH 11.4) groundwater environment (Tiago et al., 2006). We have also
302 detected the presence of Archaea (including methanogens) and sulfate-reducing
303 prokaryotes in the mats and from well water from Adobe Springs by PCR with rRNA
304 gene and *dsrAB* gene primer sets, though these organisms have not yet been identified via
305 sequence analysis. Many of the methanogens, detected with archaeal 16S rRNA gene
306 primers and with *mcrA* gene primers, are closely related to the *Methanobacterium*
307 *alcaliphilum* strain DSM3387, an alkaliphilic hydrogen-consuming (H₂/CO₂) methanogen
308 from a deep coal seam groundwater sample with a pH of ~8.4 (**DO649335**). The putative
309 identification of alkaliphilic organisms in the Del Puerto Creek and cultivation analyses
310 of cyanobacteria from the microbial mats suggest that the elevated pH in this
311 environment most likely exerts a selective influence on the composition of the microbial
312 communities.

313

314 **4. Discussion**

315

316 There is an extensive literature demonstrating that the presence and activity of microbial
317 populations are critical to the precipitation of carbonates, particularly magnesium-rich
318 carbonates, such as dolomite (e.g., Vasconcelos et al., 1995; Wright, 1999; Warthmann et
319 al., 2000; Barton et al., 2001; van Lith et al., 2003; Roberts et al., 2004; Altermann et al.,
320 2006). Microbial involvement in carbonate precipitation has been demonstrated for
321 stratified, laminated structures such as stromatolites (Dupraz and Visscher, 2005), and
322 these structures, generally composed of limestone or dolomite, have been found in the

323 sedimentary record dating back almost 3.5 billion years (Awramik, 1984; Altermann et
324 al. 2006). The best-studied environments for production of stromatolites are marine or
325 hypersaline environments. Although such systems have relatively high concentrations of
326 sulfate, which generally inhibits the precipitation of dolomite (Baker and Kastner, 1981),
327 dolomite or Mg-rich carbonates can be precipitated under appropriate environmental
328 conditions. Microorganisms can provide the conditions required for precipitation of
329 carbonates: elevated pH (photosynthesis and anaerobic respiration), elevated dissolved
330 inorganic carbon (respiration), and nucleation sites from extracellular polymeric
331 substances (EPS), or degradation of EPS resulting in the release of cations (Dupraz and
332 Visscher, 2005). However, microbial activities may also inhibit the precipitation of
333 carbonates, by cation capture by EPS, consumption of DIC, and acidification (sulfide
334 oxidation) (Barron et al., 2006; Dupraz et al., 2004; Dupraz and Visscher, 2005; Hartley
335 et al., 1996). In marine environments, the key microbial functions involved in the
336 precipitation of carbonates appear to be photosynthesis and anaerobic heterotrophic
337 oxidation of organic matter, generally coupled to sulfate reduction (cf., Visscher et al.,
338 1998; Wright and Altermann, 2000; Visscher et al., 2000; Visscher and Stolz, 2005;
339 Altermann et al., 2006).

340

341 In alkaline, hypersaline lakes in South Australia, the heightened activity of sulfate-
342 reducing bacteria (SRB) during seasonal evaporation events was correlated with the
343 precipitation of dolomite (Wright, 1999). Carbonate deposits can also occur under
344 freshwater conditions, and have been observed in association with alkaline springs
345 emanating from altered ophiolites (Barnes and O'Neil, 1971). While cyanobacterial

346 activity has been implicated for some freshwater carbonate deposits (e.g. Freytet and
347 Verrecchia, 1998; Merz-Preiss and Riding, 1999), the association of microbial activity
348 with carbonates precipitating in ophiolite environments has not been studied in detail.
349 However, in our initial characterization of the Adobe Creek locality, identified
350 populations of alkaliphilic organisms in the Del Puerto Creek and cyanobacteria from the
351 microbial mats are similar in nature to the types of organisms encountered in stromatolite
352 ecosystems, which are closely linked to biological precipitation of carbonates. The good
353 match between $\delta^{18}\text{O}$ values calculated using the microbially mediated isotopic
354 fractionation equation of Vasconcelos (2005) and measured $\delta^{18}\text{O}$ values from the
355 laminated carbonates supports the idea that precipitation of dolomites at Adobe Springs
356 under ambient temperature conditions (18–24°C) is facilitated by the presence of the
357 alkaliphilic microbial community.

358

359 **5. Conclusions**

360

361 The process of serpentinization of mafic and ultramafic rocks produces Mg-rich alkaline
362 waters, which are associated with Mg-Ca carbonate cements and unusual microbial
363 communities. The process of serpentinization can generate methane and hydrogen, two
364 potential sources of energy for chemosynthetic organisms. Such a setting (where water is
365 in contact with mafic and ultramafic rocks) may serve as a good analog for similar
366 environments on Mars that may be capable of supporting life.

367

368 We have focused our initial investigation on three critical components of the Adobe
369 Springs system: 1) the chemistry of the alkaline waters emanating from mafic and
370 ultramafic rocks; 2) the types and compositions of actively precipitating carbonate
371 cements found lining the adjacent creek drainages, and; 3) the novel microbial
372 communities associated with the alkaline waters and carbonate cements. The deposition
373 of dolomite cements from these low temperature cements may require microbial
374 mediation, which would thus represent a biosignature of this particular biogeochemical
375 environment.

376

377 Additional work is needed to confirm the hypothesis that serpentinite-associated
378 carbonate cements can be a biosignature. One possible approach would be to examine
379 the stable isotope composition of carbon in the cements to ascertain whether they contain
380 a biogenic signature (e.g., García del Cura et al., 2001; Peckman et al., 1999; Cavagna et
381 al., 1999). Laboratory precipitation experiments conducted using sterilized stream fluids
382 with and without microbial cultures selected from those identified in the alkaline waters
383 may also provide information on the possible role that biomineralization may play in the
384 generation of the carbonate cements, in particular, the dolomite. If such a link can be
385 demonstrated, then dolomite precipitation in hydrothermally altered mafic and ultramafic
386 rocks could be used as a biomarker on Mars and other planets.

387

388 **Acknowledgments**

389 Financial support for our work at Adobe Springs came from the NASA Astrobiology
390 Institute Grant (“Linking Our Origins to Our Future”, P.I. David Des Marais,

391 NASA/Ames Research Center) and a sub-contract to the SETI Institute (Cooperative
392 Agreement NNA06CB35A). Additional financial support came from the NASA
393 Postdoctoral Program, managed by Oak Ridge Associated Universities. Support to P.
394 Dobson at Lawrence Berkeley National Laboratory was provided under Contract No. DE-
395 AC02-05CH11231 with the U.S. Department of Energy. Wisc-SIMS, the Wisconsin
396 SIMS Laboratory, is partially funded by NSF-EAR (0319230, 0509639, 0744079), DOE
397 (93ER14389), and the NASA Astrobiology Institute.

398

399 This work was a natural outgrowth of a related project conducted by Professor Mitch
400 Schulte (U. Missouri) and Dr. Dave Blake (NASA/Ames) and funded by NASA's
401 Exobiology Program. We thank Dr. Robert Coleman for his suggestion to explore the site
402 at Adobe Springs and for sharing his knowledge of the Del Puerto Ophiolite. We thank
403 Bill Evans and Bob Mariner and other members of the Hydrology Branch of the U.S.
404 Geological Survey, Menlo Park, CA, for discussions concerning California spring
405 chemistry. Bob Mariner shared the field notebook of the late Ivan Barnes, who studied
406 the waters at Adobe Springs extensively in the 1960's. We thank our colleagues Kendra
407 Turk and Mike Kubo (SETI Institute) and Alaina Brinley (National Science Foundation
408 Research Experience for Undergraduates Grant to the SETI Institute, P.I. Cynthia
409 Phillips) for laboratory and field assistance. We appreciate the assistance of Linda L.
410 Jahnke in visual characterization of cyanobacterial isolates from the microbial mats. We
411 also wish to thank Paul Mason (Mgwaters.com) for granting permission to conduct this
412 work on his property and for his enthusiastic support of this project. We thank Gian

413 Gabriele Ori, Goro Komatsu, and an anonymous reviewer for their constructive reviews
414 of this paper.

415

416 **References**

417

418 Altermann, W., Kazmierczak, J., Oren, A., and Wright, D.T., 2006. Cyanobacterial
419 calcification and its rock-building potential during 3.5 billion years of Earth history.
420 *Geobiology* 4, 147–166.

421 Andrews-Hanna, J.C., Phillips, R.J., and Zuber, M.T., 2007. Meridiani Planum and the
422 global hydrology of Mars. *Nature* 446, 163–166.

423 Awramik, S.M., 1984. Ancient Stromatolites and Microbial Mats, *In: Microbial Mats:*
424 *Stromatolites.* (eds. Y. Cohen, R.W. Castenholz, and H. Halvorson) New York: Alan
425 R. Liss, Inc., pp.1–22.

426 Baker, P.A., and Kastner, M., 1981. Constraints on the formation of sedimentary
427 dolomite. *Science* 213, 214–216.

428 Baker, V.R., 2006. Geomorphological Evidence for Water on Mars. *Elements* 2, 139–
429 143.

430 Barnes, I., LaMarche, Jr., V.C., and Himmelberg, G., 1967. Geochemical evidence of
431 present-day serpentinization. *Science* 156, 830–832.

432 Barnes, I., and O’Neil, J.R., 1971. Calcium-magnesium carbonate solid solutions from
433 Holocene conglomerate cements and travertines in the Coast Range of California.
434 *Geochim. Cosmochim. Acta* 35, 699–718.

435 Barron, C., Duarte, C.M., Frankignoulle, M., and Borges, A.V., 2006. Organic carbon
436 metabolism and carbonate dynamics in a Mediterranean seagrass (*Posidonia*
437 *oceanica*) meadow. *Estuaries and Coasts* 29, 417–426.

438 Barton, H.A., Spear, J.R., and Pace, N.R., 2001. Microbial life in the underworld:
439 Biogenicity in secondary mineral formations. *Geomicrobiology J.* 18, 359–368.

440 Beaty, D.W., Clifford, S.M., Borg, L.E., Catling, D.C., Craddock, R.A., Des Marais, D.J.,
441 Farmer, J.D., Frey, H.V., Haberle, R.M., McKay, C.P., Newsom, H.E., Parker, T.J.,
442 Segura, T., and Tanaka, K.L., 2005. Key science questions from the Second
443 Conference on Early Mars: Geologic, hydrologic, and climatic evolution and the
444 implications for life. *Astrobiology* 5, 663–689.

445 Blake, D.F., and Peacor, D.R., 1985. TEM/STEM microanalysis of Holocene fresh-water
446 magnesian carbonate cements from the Coast Range of California. *Amer. Mineral.*
447 70, 388–394.

448 Blank, J.G., Blake, D.F., Green, S.J., Brinley, A.I., Jahnke, L.L., Kubo, M.D., Hoehler,
449 T.M., and Des Marais, D.J., 2006. Biogeochemistry of Ca-Mg carbonate cements
450 associated with ophiolite-hosted cold springs, Coast Range, California, USA. *Geol.*
451 *Soc. Amer. Abstracts with Programs* 38, 505.

452 Blank, J.G., Valley, J.W., Treiman, A.H., Kita, N., and Blake, D.F., 2007. Oxygen
453 isotope variation in Ca-Mg carbonate cements in the California Coast Range
454 Ophiolite: Geochemistry of Martian analog environments. *Lunar Planet. Sci. Conf.*
455 38, Abstract 2150.

456 Boston, P.J., Ivanov, M.V., and Mckay, C.P., 1992. On the possibility of chemosynthetic
457 ecosystems in subsurface habitats on Mars. *Icarus* 95, 300–308.

458 Bowman, J.R., Valley, J.W., and Kita, N.T., 2008. Mechanisms of oxygen isotopic
459 exchange and isotopic evolution of $^{18}\text{O}/^{16}\text{O}$ -depleted periclase zone marbles in the
460 Alta aureole, Utah—Insights from ion microprobe analysis of calcite. *Contrib. Min.
461 Pet.* (in press).

462 Casamayor, E.O., Massana, R., Benlloch, S., Øvreås, L., Díez, B., Goddard, V.J., Gasol,
463 J.M., Joint, I., Rodríguez-Valera, F., and Pedrós-Alió, C., 2002. Changes in archaeal,
464 bacterial and eukaryal assemblages along a salinity gradient by comparison of genetic
465 fingerprinting methods in a multipond solar saltern. *Environ. Microbiol.* 4, 338–348.

466 Cavagna, S., Clari, P., and Martire, L., 1999. The role of bacteria in the formation of cold
467 seep carbonates: geological evidence from Monferrato (Tertiary, NW Italy). *Sed.
468 Geol.* 126, 253–270.

469 Christensen, P.R., McSween, H.Y., Jr., Bandfield, J.L., Ruff, S.W., Rogers, A.D.,
470 Hamilton, V.E., Gorelick, N., Wyatt, M.B., Jakosky, B.M., Kieffer, H. H., Malin,
471 M.C., and Moersch, J.E., 2005. Evidence for magmatic evolution and diversity on
472 Mars from infrared observations. *Nature* 436, 504–509.

473 Dietrich, W.E. and Perron, J.T., 2006. The search for a topographic signature of life.
474 *Nature* 439, 411-419.

475 Dupraz, C., and Visscher, P.T., 2005. Microbial lithification in marine stromatolites and
476 hypersaline mats. *Trends in Microbiology* 13, 429–438.

477 Dupraz, C., Visscher, P.T., Baumgartner, L.K., and Reid, R.P., 2004. Microbe-mineral
478 interactions: early carbonate precipitation in a hypersaline lake (Eluthera Island,
479 Bahamas). *Sedimentology* 51, 745–765.

480 Eiler, J.M., Valley, J.W., Graham, C.M., and Fournelle, J., 2002. Two populations of
481 carbonate in ALH84001: Geochemical evidence for discrimination and genesis.
482 *Geochim. Cosmochim. Acta* 66, 1285-1303.

483 Evarts, R.C., 1977. The geology and petrology of the Del Puerto Ophiolite, Diablo
484 Range, central California Coast Ranges. In: *North American ophiolites* (eds. R.G.
485 Coleman and W.P. Irwin). Oregon Dept. Geol. Mineral Ind. Bull. 95, 121–139.

486 Evarts, R.C., and Schiffman, P., 1983. Submarine hydrothermal alteration of the Del
487 Puerto Ophiolite, California. *Amer. J. Sci.* 283, 289–340.

488 Evarts, R.C., Sharp, W.D., and Phelps, D.W., 1992. The Del Puerto Canyon remnant of
489 the Great Valley Ophiolite: Geochemical and age constraints on its formation and
490 evolution. *Amer. Assoc Petrol. Geol. Bull.* 76, 418.

491 Fisk, M.R., and Giovannoni, S.J., 1999. Sources of nutrients and energy for a deep
492 biosphere on Mars. *J. Geophys. Res.* 104, 11805–11816.

493 Formisano, V., Atreya, S., Encrenaz, T., Ignatiev, N., and Giuranna, M., 2004. Detection
494 of methane in the atmosphere of Mars. *Science* 306, 1758–1761.

495 Freytet, P., and Verrecchia, E.P., 1998. Freshwater organisms that build stromatolites: a
496 synopsis of biocrystallization by prokaryotic and eukaryotic algae. *Sedimentology* 45,
497 535–563.

498 Friedman, I., and O’Neil, J.R., 1977. Compilation of stable isotope fractionation factors
499 of geochemical interest. USGS Prof. Paper 440-KK.

500 Früh-Green, G.L., Kelley, D.S., Bernasconi, S.M., Karston, J.A., Ludwig, K.A.,
501 Butterfield, D.A., Boschi, C., and Proskurowski, G., 2003. 30,000 years of
502 hydrothermal activity at the Lost City Vent Field. *Science* 301, 495–498.

503 García del Cura, M.A., Calvo, J.P., Ordóñez, S., Jones, B.F., and Cañaveras, J.C., 2001.
504 Petrographic and geochemical evidence for the formation of primary, bacterially
505 induced lacustrine dolomite: La Roda "white earth" (Pliocene, central Spain).
506 *Sedimentology* 48, 897–915.

507 Geets, J., Borremans, B., Diels, L., Springael, D., Vangronsveld, J., Van der Lelie, D.,
508 and Vanbroekhoven, K., 2006. *DsrB* gene-based DGGE for community and diversity
509 surveys of sulfate-reducing bacteria. *J. Microbiol. Methods* 66, 194–205.

510 Green, S.J., Blackford, C., Bucki, P., Jahnke, L.L., Bebout, B.M., and Prufert-Bebout, L.,
511 2008. A salinity and sulfate manipulation of hypersaline microbial mats reveals stasis
512 in the cyanobacterial community structure. *ISME Journal* 2, 457–470.

513 Hartley, A.M., House, W.A., Leadbeater, B.S.C., and Callow, M.E., 1996. The use of
514 microelectrodes to study the precipitation of calcite upon algal biofilms. *J. Colloid*
515 *Interface Sci.* 183, 498–505.

516 Horita, J., and Clayton, R.N., 2007. Comment on the studies of oxygen isotope
517 fractionation between calcium carbonates and water at low temperatures by Zhou and
518 Zheng (2003; 2005). *Geochim. Cosmochim. Acta* 71, 3131–3135.

519 Kelley, D.S., Karson, J.A., Früh-Green, G.L., Yoerger, D.R., Shank, T.M., Butterfield,
520 D.A., Hayes, J.M., Schrenk, M.O., Olson, E.J., Proskurowski, G., Jakuba, M.,
521 Bradley, A., Larson, B., Ludwig, K., Glickson, D., Buckman, K., Bradley, A.S.,
522 Brazelton, W.J., Roe, K., Elend, M.J., Delacour, A., Bernasconi, S.M., Lilley, M.D.,
523 Baross, J.A., Summons, R.E., and Sylva, S.P., 2005. A Serpentinite-Hosted
524 Ecosystem: The Lost City Hydrothermal Field. *Science* 307, 1428–1434.

525 Kita N.T., Ushikubo, T., Fu, B., Spicuzza M.J., and Valley, J.W., 2007. Analytical
526 developments on oxygen three isotope analyses using a new generation ion
527 microprobe IMS-1280. Lunar Planet. Sci. Conf. 38, Abstract 1981.

528 Knoll, A.H., and Grotzinger, J., 2006. Water on Mars and the prospect of Martian life.
529 Elements 2, 169–173.

530 Land, L.S., 1998. Failure to precipitate dolomite at 25°C from dilute solution despite
531 1000-fold oversaturation after 32 years. Aquatic Geochem. 4, 361–368.

532 Leshin, L., McKeegan, K., and Harvey, R., 1998. Oxygen isotopic constraints on the
533 genesis of carbonates from Martian meteorite ALH84001. Geochim. Cosmochim.
534 Acta 62, 3–13.

535 Leshin, L.A., and Vicenzi, E., 2006. Aqueous processes recorded by Martian meteorites:
536 Analyzing Martian water on Earth. Elements 2, 157–162.

537 Luton, P.E., Wayne, J.M., Sharp, R.J., and Riley, P.W., 2002. The mcrA gene as an
538 alternative to 16S rRNA in the phylogenetic analysis of methanogen populations in
539 landfill. Microbiology 148, 3521–3530.

540 McEwen, A.S. , Hansen, C.J., Delamere, W.A., Eliason, E.M., Herkenhoff, K.E.,
541 Keszthelyi, L., Gulick, V.C., Kirk, R.L., Mellon, M.T., Grant, J.A., Thomas N.,
542 Weitz, C.M. Squyres, S.W., Bridges, N.T., Murchie, S.L., Seelos, F., Seelos, K.,
543 Okubo, C.H., Milazzo, M.P., Tornabene, L.L., Jaeger, W.L., Byrne, S., Russell, P.S.,
544 Griffes, J.L., Martínez-Alonso, S., Davatzes, A., Chuang, F.C., Thomson, B.J.,
545 Fishbaugh, K.E., Dundas C.M., Kolb, K.J., Banks, M.E., and Wray, J.J., 2007. A
546 Closer Look at Water-Related Geologic Activity on Mars. Science 317, 1706–1709.

547 Merz-Preiss, M., and Riding, R., 1999. Cyanobacterial tufa calcification in two
548 freshwater streams: ambient environment, chemical thresholds and biological
549 processes. *Sed. Geol.* 126, 103–241.

550 Ming, D.W., Mittlefehldt D.W., Morris, R.V. , Golden, D.C., Gellert, R., Yen, A., Clark,
551 B.C., Squyres, S.W., Farrand, W.H., Ruff, S.W., Arvidson, R.E., Klingelhöfer, G.,
552 McSween, H.Y., Rodionov, D.S., Schröder, C., de Souza Jr., P.A., and Wang, A..
553 2006, Geochemical and mineralogical indicators for aqueous processes in the
554 Columbia Hills of Gusev crater, Mars. *J. Geophys. Res.* 111, E02S12,
555 doi:10.1029/2005JE002560

556 Muyzer, G., de Waal, E.C., and Uitterlinden, A.G., 1993. Profiling of complex microbial-
557 populations by denaturing gradient gel-electrophoresis analysis of polymerase chain
558 reaction-amplified genes-coding for 16S Ribosomal-RNA. *Appl. Environ. Microbiol.*
559 59, 695–700.

560 Muyzer, G., and Smalla, K., 1998. Application of denaturing gradient gel electrophoresis
561 (DGGE) and temperature gradient gel electrophoresis (TGGE) in microbial ecology.
562 *Antonie van Leeuwenhoek* 73, 127–141.

563 Newsom, H.E., Shearer, C.K., and Treiman, A.H., 2001. Mobile elements determined by
564 SIMS analysis in hydrous alteration materials in the Lafayette Martian meteorite.
565 *Lunar Planet. Sci. Conf.* 32, Abstract 1396.

566 Nubel, U., Garcia-Pichel F., and Muyzer, G., 1997. PCR primers to amplify 16S rRNA
567 genes from cyanobacteria. *Appl. Environ. Microbiol.* 63, 3327–3332.

568 O'Neil, J.R., Clayton, R.N., and Mayeda, T.K., 1969. Oxygen isotope fractionation in
569 divalent metal carbonates. *J. Chem. Phys.* 51, 5547–5558.

570 O'Neil, J.R., and Barnes, I., 1971. C¹³ and O¹⁸ compositions in some fresh-water
571 carbonates associated with ultramafic rocks and serpentinites: western United States.
572 *Geochim. Cosmochim. Acta* 35, 687–697.

573 Page F.Z., Ushikubo, T., Kita, N.T., Riciputi, L.R., and Valley, J.W., 2007. High-
574 precision oxygen isotope analysis of picogram samples reveals 2 μm gradients and
575 slow diffusion in zircon. *Amer. Mineral.* 92, 1772–1775.

576 Peckman, J., Theil, V., Michaelis, W., Clari, P., Gaillard, C., Martire, L., and Reitner, J.,
577 1999. Cold seep deposits of Beauvoisin (Oxfordian; southeastern France) and
578 Marmorito (Miocene; northern Italy): microbially induced authigenic carbonates. *Int.*
579 *J. Earth Sci.* 88, 60–75.

580 Pentecost, A., 2005. *Travertine*. Berlin, Springer-Verlag.

581 Reed, M., 1982. Calculation of multicomponent chemical equilibria and reaction
582 processes in systems involving minerals, gases and an aqueous phase. *Geochim.*
583 *Cosmochim. Acta* 46, 513–528.

584 Roberts, J.A., Bennett, P.C., González, L.A., Macpherson, G.L., and Milliken, K.L.,
585 2004. Microbial precipitation of dolomite in methanogenic groundwater. *Geology* 32,
586 277–280.

587 Schmidt, M., Xeflide, S., Botz, R., and Mann, S., 2005. Oxygen isotope fractionation
588 during synthesis of CaMg-carbonate and implications for sedimentary dolomite
589 formation. *Geochim. Cosmochim. Acta* 69, 4665–4674.

590 Schulte, M., Blake, D., Hoehler, T., and McCollom, T., 2006. Serpentinization and its
591 implications for life on the early Earth and Mars. *Astrobiology* 6, 364–376.

592 Shervais, J.W., Kolesar, P., and Andreasen, K., 2005a. A field and chemical study of
593 serpentinization – Stonyford, California: Chemical fluxes and mass balance. *Int.*
594 *Geol. Rev.* 47, 1–23.

595 Shervais, J.W., Murchey, B.L., Kimbrough, D.L., Renne, P.R., and Hanan, B., 2005b.
596 Radioisotopic and biostratigraphic age relations in the Coast Range Ophiolite,
597 northern California: Implications for the tectonic evolution of the Western Cordillera.
598 *Geol. Soc. Amer. Bull.* 117, 633–653.

599 Singer, R.B., and McSween, H.Y., 1993. The igneous crust of Mars: compositional
600 evidence from remote sensing and the SNC Meteorites, *in: Resources of near-Earth*
601 *Space* (J.S. Lewis, M.S. Matthews, and M.L. Guerrieri, eds.), University of Arizona
602 Press, Tucson AZ, pp. 709–736.

603 Sleep N.H., Meibom, A., Fridriksson, Th., Coleman, R.G., and Bird, D.K., 2004. H₂-rich
604 fluids from serpentinization: Geochemical and biotic implications. *Proc. Nat. Acad.*
605 *Sci.* 101, 12818–12823.

606 Squyres, S.W., Arvidson, R.E., Bell, III, J.F., Brückner, J., Cabrol, N.A., Calvin, W.,
607 Carr, M.H., Christensen, P.R., Clark, B.C., Crumpler, L., Des Marais, D.J., d'Uston,
608 C., Economou, T., Farmer, J., Farrand, W., Folkner, W., Golombek, M., Gorevan, S.,
609 Grant, J.A., Greeley, R., Grotzinger, J., Haskin, L., Herkenhoff, K.E., Hviid, S.,
610 Johnson, J., Klingelhöfer, G., Knoll, A.H., Landis, G., Lemmon, M., Li, R., Madsen,
611 M.B., Malin, M.C., McLennan, S.M., McSween, H.Y., Ming, D.W., Moersch, J.,
612 Morris, R.V., Parker, T., Rice, Jr., J.W., Richter, L., Rieder, R., Sims, M., Smith, M.,
613 Smith, P., Soderblom, L.A., Sullivan, R., Wänke, H., Wdowiak, T., Wolff, M., and

614 Yen, A., 2004a. The Opportunity Rover's Athena Science Investigation at Meridiani
615 Planum, Mars. *Science* 306, 1698–1703.

616 Squyres, S.W., Grotzinger, J.P., Arvidson, R.E., Bell, III, J.F., Calvin, W. Christensen,
617 P.R., Clark, B.C., Crisp, J.A., Farrand, W.H., Herkenhoff, K.E., Johnson, J.R.,
618 Klingelhöfer, G., Knoll, A.H., McLennan, S.M., McSween, Jr., H.Y., Morris, R.V.,
619 Rice, Jr., J.W., Rieder, R., and Soderblom, L.A., 2004b. *In situ* evidence for an
620 ancient aqueous environment at Meridiani Planum. *Mars. Science* 306, 1709–1714.

621 Squyres, S.W., Knoll, A.H., Arvidson, R.E., Clark, B.C., Grotzinger, J.P., Jolliff, B.L.,
622 McLennan, S.M., Tosca, N., Bell, III, J.F., Calvin, W.M., Farrand, W.H., Glotch,
623 T.D., Golombek, M.P., Herkenhoff, K.E., Johnson, J.R., Klingelhöfer, G., McSween,
624 H.Y., and Yen, A.S., 2006. Two Years at Meridiani Planum: Results from the
625 Opportunity Rover. *Science* 313, 1403–1407.

626 Squyres, S.W. and Knoll, A.H., 2005. Sedimentary rocks at Meridiani Planum: Origin,
627 diagenesis, and implications for life on Mars. *Earth Planet. Sci. Lett.* 240, 1–10.

628 Surour, A.A., and Arafa, E.H., 1997. Ophicarbonates: calichified serpentinites from
629 Gebel Mohagara, Wadi Ghadir area, Eastern Desert, Egypt. *J. African Earth Sci.* 24,
630 315–324.

631 Tarutani, T., Clayton, R.N., and Mayeda, T.K., 1969. The effect of polymorphism and
632 magnesium substitution on oxygen isotope fractionation between calcium carbonates
633 and water. *Geochim. Cosmochim. Acta* 33, 987–996.

634 Tiago, I., Mendes, V., Pires, C., Morais, P.V., and Verissimo, A., 2006. *Chimaereicella*
635 *alkaliphila* gen. nov., sp. nov., a Gram-negative alkaliphilic bacterium isolated from a
636 nonsaline alkaline groundwater. *Syst. Appl. Microbiol.* 29, 100–108.

637 Treiman A.H., 1998. The history of Allan Hills 84001 revised: Multiple shock events.
638 Meteoritics Planet. Sci. 33, 753–764.

639 Treiman, A.H., and Goodrich, C.A., 2002. Pre-terrestrial aqueous alteration of the
640 Y000593 and Y000749 Nakhlite meteorites. Nat. Inst. Polar Res. Symp. Antarctic
641 Meteorites XXVII, 166–167.

642 Treiman A.H. and Romanek, C.S., 1998. Bulk and stable isotopic compositions of
643 carbonate minerals in Martian meteorite Allan Hills 84001: No proof of high
644 formation temperature. Meteoritics Planet. Sci. 33, 737–742.

645 Treiman A.H., Lanzirotti, A., and Xirouchakis, D., 2004. Ancient water on asteroid 4
646 Vesta: Evidence from a quartz veinlet in the Serra de Magé eucrite meteorite. Earth
647 Planet. Sci. Lett. 219, 189–199.

648 Valley, J.W., Eiler, J.M., Graham, C.M., Gibson, E.K., Romanek, C.S., and Stolper,
649 E.M., 1997. Low-temperature carbonate concretions in the Martian meteorite
650 ALH84001: Evidence from stable isotopes and mineralogy. Science 275, 1633–1668.

651 Van Lith, Y., Warthmann, R., Vasconcelos, C., and McKenzie, J.A., 2003. Microbial
652 fossilization in carbonate sediments: a result of the bacterial surface involvement in
653 dolomite precipitation. Sedimentology 50, 237–245.

654 Vasconcelos, C., McKenzie, J.A., Bernasconi, S., Grujic, D., and Tien, A.J., 1995.
655 Microbial mediation as a possible mechanism for natural dolomite formation at low
656 temperatures. Nature 377, 220–222.

657 Vasconcelos, C., McKenzie, J.A., Warthmann, R., and Bernasconi, S.M., 2005.
658 Calibration of the $\delta^{18}\text{O}$ paleothermometer for dolomite precipitated in microbial
659 cultures and natural environments. Geology 33, 317–320.

660 Visscher, P.T., Reid, R.P., and Bebout, B.M., 2000. Microscale observations of sulfate
661 reduction: correlation of microbial activity with lithified micritic laminae in modern
662 marine stromatolites. *Geology* 28, 919–922.

663 Visscher, P.T., Reid, R.P., Bebout, B.M., Hoefft, S.E., Macintyre, I.G., and Thompson,
664 Jr., J.A., 1998. Formation of lithified micritic laminae in modern marine stromatolites
665 (Bahamas): The role of sulphur cycling. *Amer. Mineral.* 83, 1482–1493.

666 Visscher, P.T., and Stolz, J.F., 2005. Microbial mats as bioreactors: populations,
667 processes, and products. *Palaeogeog., Palaeoclimatol., Palaeoecol.* 219, 87–100.

668 Warthmann, R., van Lith, Y., Vasconcelos, C., McKenzie, J.A., and Karpoff, A.M., 2000.
669 Bacterially induced dolomite precipitation in anoxic culture experiments. *Geology* 28,
670 1091–1094.

671 Wright, D.T., 1999. The role of sulfate-reducing bacteria and cyanobacteria in dolomite
672 formation in distal ephemeral lakes of the Coorong region, South Australia. *Sed.*
673 *Geol.* 126, 147–157.

674 Wright, D.T., and Altermann, W., 2000. Microfacies development in Late Archaean
675 stromatolites and oolites of the Campbellrand Subgroup, South Africa. *In: Carbonate*
676 *Platform Systems. Components and Interactions* (Insalco, E., P.W. Skelton, and T.J.
677 Palmer, eds.). *Geol.Soc. London Spec. Pub.* 178, 51–70.

678 Wyatt, M.B., and McSween, Jr., H.Y., 2006. The orbital search for altered materials on
679 Mars. *Elements* 2, 145–150.

680 **Figure Captions**

681

682 **Figure 1.** Field site showing locations of the three sampling sites (indicated by the push
683 pin icons) associated with alkaline waters in the Del Puerto Ophiolite, CA: **DP6**, at the
684 Del Puerto Creek, **Adobe Springs Well**, and **AC6**, at Adobe Creek, a tributary of Del
685 Puerto Creek. Figure made using GoogleEarth.

686

687 **Figure 2. (A-H)** Photographs of carbonate cements and microbial communities from the
688 Adobe Springs sampling sites, April-June 2006. **(A)** Del Puerto Creek (DPC) and **(B)**
689 Adobe Creek, June 2006, showing carbonate cements and microbial biomass
690 (periphyton). **(C)** Hand sample of DPC carbonate cement. **(D)** DPC, 10 miles
691 downstream of sampling site, parallel to the year-round main creek flow. **(E)** Thin section
692 of serpentine grain bordered by banded carbonate from DPC. White scale bar indicates 1
693 mm (horizontal and vertical). **(F)** Thin, laminated microbial mat underlain by anaerobic
694 mud (AC6). **(G)** *Leptolyngbya*-like and **(H)** *Arthrospira*-like cyanobacteria recovered
695 from microbial mat samples. Scale bars indicate 30 μm .

696

697 **Figure 3.** Mole fraction of major cations of carbonate cement from Del Puerto Creek, as
698 determined from electron microprobe analysis.

699

700 **Figure 4.** Results of EMP and SIMS analysis of a banded cement from the Del Puerto
701 Creek. **(A)** Photomicrograph of sample in transmitted light, illustrating fine-scale Mg-
702 Ca carbonate laminae deposited outward from a serpentinized clast. In-situ oxygen

703 isotopic measurements were made using a CAMECA ims-1280 SIMS at the University
704 of Wisconsin; transect points (in white) were created by the SIMS beam. The polished
705 sample surface was coated with a thin layer of gold prior to analysis; gold in and
706 adjacent to the analysis pits was sputtered during analysis, leaving gold-free regions
707 wider than their corresponding pits (here, the pits are ~8 or ~15 microns diameter) in the
708 sample. Yellow scale bar represents 100 μm ; width of cement section is ~ 550 μm . (B)
709 Variation in $\delta^{18}\text{O}$ and Ca# (the mole fraction of $\text{Ca}/(\text{Ca} + \text{Mg})$) as a function of distance
710 from the serpentine grain boundary.



DP 6

N

Del Puerto Creek

Adobe Creek

Adobe Springs Well

AC 6

500 m

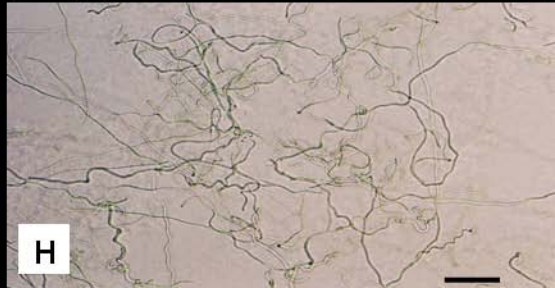
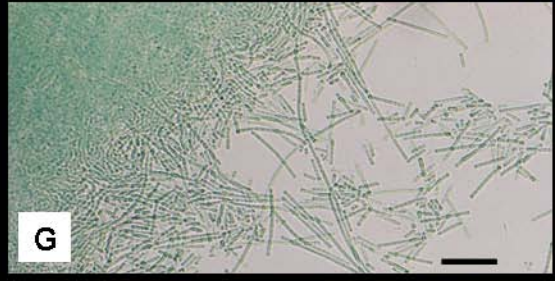
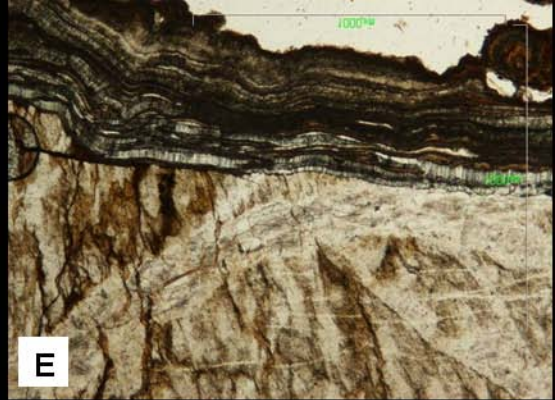
Image © 2008 DigitalGlobe

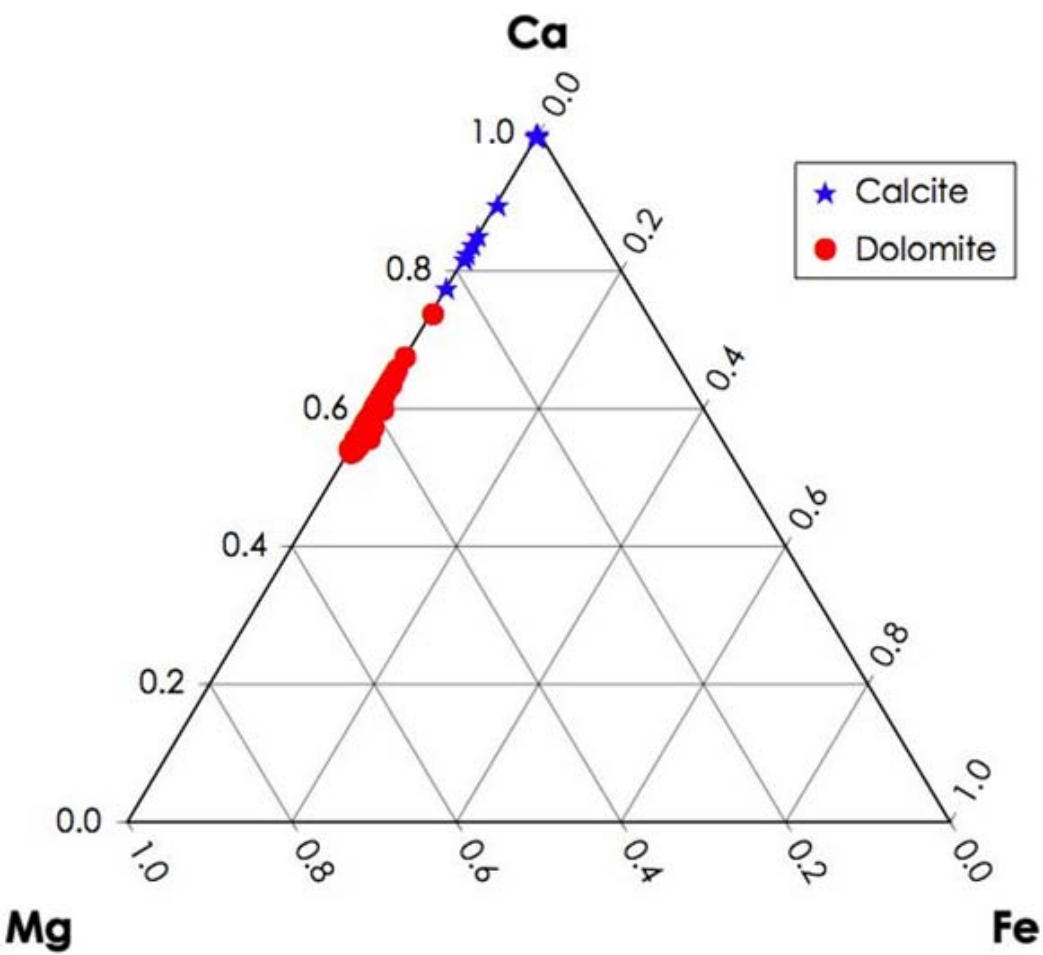
Google

Nov 23, 2004

Eye alt 1.61 km

37°24'24.96" N 121°24'34.12" W





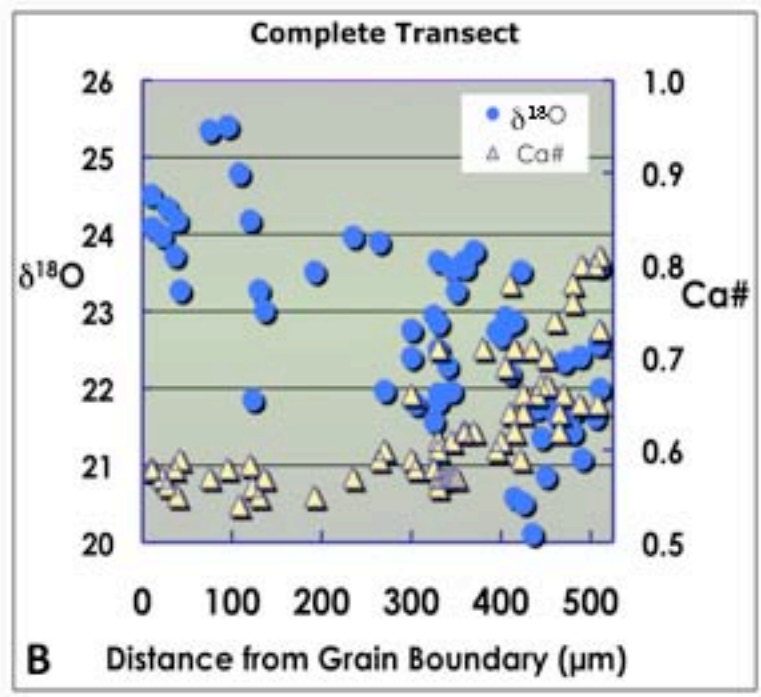
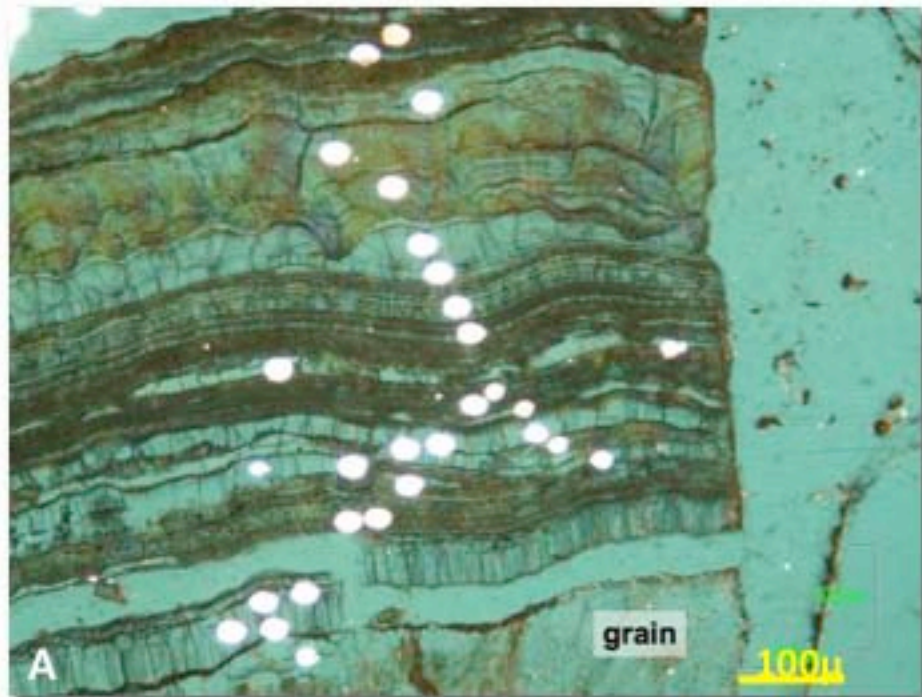


Table 1: Fluid chemistry of representative water samples

	Adobe Springs Well	Del Puerto Creek
Collection date	June 9, 2007	June 9, 2007
Ca ²⁺	3.5	8.1
Mg ²⁺	110	150
K ⁺	0.31	0.6
Na ⁺	5.4	9.6
HCO ₃ ⁻	400	550
CO ₃ ²⁻	66	89
Cl ⁻	4.8	9.5
NH ₃ (total as N)	0.01	0.018
SO ₄ ²⁻	16	10
SiO ₂	5.6	13
OH ⁻	<1.6	<1.6
Alkalinity as CaCO ₃	440	600
Field pH	8.73	8.52
Lab pH	8.69	8.61
Collection T°C	17.8	24.2
δ ¹⁸ O	-7.9	-7.1
δD	-57	-52

Concentrations of dissolved species given in mg/L; isotopic values reported in permil relative to VSMOW.

Table 2: Calculated Equilibrium Carbonate Oxygen Isotope Compositions

	Adobe Springs Well	Del Puerto Creek
Collection date	June 9, 2007	June 9, 2007
Collection T°C	17.8	24.2
$\delta^{18}\text{O}_{\text{VSMOW}}$ (per mil)	-7.9	-7.1
Predicted dolomite compositions (‰)		
Tarutani et al. (1969)	25.3	24.7
Schmidt et al. (2005)	26.6	26.1
Vasconcelos et al. (2005)	24.9	24.3
Predicted calcite compositions (‰)		
O'Neil et al. (1969)	22.3	21.7
Horita and Clayton (2007)	21.8	21.3

Fractionation equations used:

$1000 \ln \alpha = 2.78 \times 10^6 T^{-2} + 0.11$ (Tarutani et al., 1969; corrected in Friedman and O'Neil, 1977, for the case of Mg mole fraction = 0.5)

$1000 \ln \alpha = 2.63 \times 10^6 T^{-2} + 3.12$ (Schmidt et al., 2005)

$1000 \ln \alpha = 2.73 \times 10^6 T^{-2} + 0.26$ (Vasconcelos et al., 2005)

$1000 \ln \alpha = 2.78 \times 10^6 T^{-2} - 2.89$ (O'Neil et al., 1969; corrected in Friedman and O'Neil, 1977)

$1000 \ln \alpha = 0.9521 \times 10^6 T^{-2} + 11.59 \times 10^6 T^{-1} - 21.56$ (Horita and Clayton, 2007)