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Spatially Variable Geothermal Heat Flux in West Antarctica: Evidence and Implications

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Abstract Geothermal heat flux (GHF) is an important part of the basal heat budget of continental ice sheets. The difficulty of measuring GHF below ice sheets has directly hindered progress in the understanding of ice sheet dynamics. We present a new GHF measurement from below the West Antarctic Ice Sheet, made in subglacial sediment near the grounding zone of the Whillans Ice Stream. The measured GHF is 88 ± 7 mW m⁻², a relatively high value compared to other continental settings and to other GHF measurements along the eastern Ross Sea of 55 mW m⁻² and 69 ± 21 mW m⁻² but within the range of regional values indicated by geophysical estimates. The new GHF measurement was made ~100 km from the only other direct GHF measurement below the ice sheet, which was considerably higher at 285 ± 80 mW m⁻², suggesting spatial variability that could be explained by shallow magmatic intrusions or the advection of heat by crustal fluids. Analytical calculations suggest that spatial variability in GHF exceeds spatial variability in the conductive heat flux through ice along the Siple Coast. Accurate GHF measurements and high-resolution GHF models may be necessary to reliably predict ice sheet evolution, including responses to ongoing and future climate change.

1. Introduction

Geothermal heat flux (GHF) is a significant source of heat in polar subglacial environments. It affects the temperature at the base of ice sheets, impacting the ice sheet mass balance directly through basal melting or freezing. GHF can have a large indirect effect on ice sheet mass balance when it brings the basal temperature above the melting point, because the presence of basal meltwater reduces basal resistance, facilitating fast sliding of ice (Weertman, 1964). GHF is prescribed as part of the lower boundary conditions for ice sheet models, which calculate patterns of basal melting and freezing to determine the degree of ice sliding. Ice sheet models are sensitive to the magnitude and spatial variability of GHF, particularly when the GHF contribution shifts basal temperatures across the melting point (Bougamont et al., 2015; Pittard et al., 2016).

Despite the importance of GHF below ice sheets, there are relatively few direct measurements of this key parameter (Davies & Davies, 2010), mainly because it is so difficult to access the subglacial environment. Prior to this study, the only direct GHF measurement below the West Antarctic Ice Sheet (WAIS) was made at Subglacial Lake Whillans (SLW) (Fisher et al., 2015); estimates were made at two additional locations using basal ice temperatures and assumptions about local ice dynamics (Clow et al., 2012; Engelhardt, 2004a). GHF has been inferred for some regions of the WAIS from the distribution of subglacial water (Schroeder et al., 2014; Siegert & Dowdeswell, 1996). Due to the paucity of observations, the GHF distribution used in ice sheet models typically falls within a relatively narrow range and has low spatial variability, based on geological or remotely sensed properties of the underlying lithosphere (An et al., 2015; Burton-Johnson et al., 2017; Fox Maule et al., 2005; Pollack et al., 1993; Shapiro & Ritzwoller, 2004). GHF models of West Antarctica are inconsistent with one another in both magnitude and distribution (Figure S1 in the supporting information), suggesting that GHF is not well constrained.

2. Materials and Methods

We determined the GHF 3 km downstream of the Whillans ice stream Grounding Zone (WGZ) using an ice borehole to collect measurements of thermal gradient and thermal conductivity.

2.1. Temperature Gradient in Sediments

The ice drilling operations are described in Tulaczyk et al. (2014). The geothermal probe used to measure the thermal gradient is the same tool used at SLW (Fisher et al., 2015). For the present study, the geothermal...
probe was deployed twice, on 15 and 18 January 2015, resulting in a horizontal distance of 3 m between measurements due to ice movement. The probe makes subsurface measurements with three autonomous sensor/logger systems, with sensor spacing of 62 cm. Autonomous sensors/loggers were calibrated before deployment with absolute accuracy of ±0.002°C (Fisher et al., 2015). The sensors/loggers were programmed just before deployment for synchronized data collection every 2 s. After data were recovered and calibration corrections were applied, we performed an additional shift to individual sensors (0.003–0.008°C) based on measurements made when the geothermal probe was held stationary in the water column (Figure S2). This is the routine approach for GHF measurements in the deep sea and assures that small variations in apparent temperature (generally due to electronic drift) do not bias geothermal data.

After the probe was inserted into the sediment at WNZ, it was held still for ~10 min to record the transient temperature response. Data from this measurement period for each sensor were fitted to a conductive heat flow model of temperature equilibration (Bullard, 1954) using TP-Fit software (Heesemann et al., 2006). The modeled equilibration period started ~100 s after penetration, to avoid deviations from the idealized model used to fit the data (a thin line source), and lasted 5–8 min. Processing of the data was managed sensor by sensor, with care taken to avoid data intervals that included evidence for probe motion, expressed as frictional heating that leads to subtle deviations in the standard equilibration curve. Data processing was completed with thermal conductivity values that are consistent with measurements described in section 2.2. Equilibration of conventional oceanographic heat flow probes often takes longer than the usual 6–7 min measurement window (Davis & Fisher, 2011), but the geothermal sensor/logger systems used in this study have sensors mounted within 5 mm outer diameter stainless steel tubing, which equilibrates quickly with surrounding material. Because of this, sensors were nearly equilibrated by the end of the useful measurement window, and extrapolation to full equilibration was relatively insensitive to model parameters (thermal conductivity, thermal diffusivity, and time shift to improve model fit). The greatest source of uncertainty in equilibrium temperature (0.001–0.006°C) came from selection of alternative measurement windows used for extrapolation to in situ conditions.

### 2.2. Thermal Conductivity

Sediment was recovered with a gravity corer in a 5.5 cm diameter polycarbonate liner through the same borehole adjacent to the thermal gradient measurements (sediment core WGZ-GC-1). Thermal conductivity, \( k \), was measured in the laboratory on a 55 cm section of this core, using the needle probe method (von Herzen & Maxwell, 1959), with measurements made every 1 cm for 40 cm. For each measurement, we drilled a 1.6 mm diameter hole through the core liner, stopping before penetrating the core itself. We placed a 5-cm-long needle probe, containing a thermistor and heater wire, through the hole and into the sediment, perpendicular to the axis of the core. Constant heating was applied, and the temperature rise during the 1 cm depth increments. Grains with diameters <1 mm were analyzed with a laser diffraction, particle size analyzer (PSA). The PSA uses light scattering to quantify particle size distribution within a liquid suspension, using a 5 mW laser source having a 750 nm wavelength. Samples were suspended in an eluent containing 0.1 g/L of sodium metaphosphate to deflocculate small particles and circulated...
continuously during measurement. The result for each sample is a probability density function of grain sizes within 93 logarithmically scaled bins ranging from \(<0.4\) μm to \(<1\) mm (Figure S4). To determine size fractions \(>1\) mm, which could not be analyzed with the PSA, samples were cut from the core and wet-sieved to isolate 1–2 mm and \(>2\) mm diameter size classes, which were weighed (Figure S5). Results from the sieve and PSA methods were combined for each sample, assuming consistent grain density in the coarse and fine fractions.

### 2.4. Spatial Variability in Other Heat Flux Terms at the Ice Sheet Bed

To place the observed GHF variations in the context of other factors influencing the basal thermal energy balance of the ice sheet, we offer basal heat flux estimates characteristic of the Siple Coast. To solve for the vertical conductive heat flux into the ice, \(q_i\), we use the analytical solution of Robin (1955) for the 1-D thermal advection-diffusion equation. This solution assumes that the vertical velocity \(v_z\) decreases linearly from the accumulation rate at the surface to 0 at the ice sheet base (Text S2 and Figure S6). We consider the steady state case for an ice sheet in mass balance to gain insight into the most important terms in the basal thermal energy balance. We take the derivative of the Robin (1955) solution, to yield the temperature gradient at the base of the ice, \(dT/dz\) \(_{ib}\), and multiply by the thermal conductivity of ice, \(k_i\), to solve for \(q_i\):

\[
q_i = k_i \frac{dT}{dz}\bigg|_{ib} = k_i \frac{2(T_b - T_s) \sqrt{P/2}}{h \sqrt{\pi} \text{ erf} \left( \sqrt{P/2} \right)}
\]

\(T_b\) and \(T_s\) are the temperature at the base and surface of the ice sheet, respectively; \(h\) is the ice thickness; and \(P\) is the Peclet number, the ratio of thermal advection to diffusion, calculated as \(a h \kappa^{-1}\), where \(a\) is the accumulation rate and \(\kappa\) is the thermal diffusivity. The parameter \(k_i\) is calculated as a function of temperature (Cuffey & Paterson, 2010).

In these calculations, we assume \(T_b\) is at the pressure melting point, \(T_m(p)\), the maximum basal temperature for a frozen bed. Thus, these \(q_i\) solutions represent a local upper bound on the vertical conductive heat flux through ice. \(T_m(p)\) is calculated using freshwater properties (Intergovernmental Oceanographic Commission et al., 2010) and \(p\) is calculated as a function of ice thickness with an average ice density of 900 kg m\(^{-3}\) to account for the effects of air bubbles and firm.

Calculated \(q_i\) values depend mainly on three independent variables: ice thickness (Fretwell et al., 2013), ice accumulation rate (Arthern et al., 2006; van de Berg et al., 2006), and mean annual surface temperature (Comiso, 2000) (error estimates in Table S6 and sensitivity analysis in Figure S7). To illustrate the contribution of variability in each of these factors to variability in \(q_i\), we present calculations of \(q_i\) along a profile near the Ross Ice Shelf grounding line, varying one factor while holding the rest at their average value across that profile (\(a\) = 12 cm yr\(^{-1}\), \(h\) = 800 m, \(T_s\) = −21°C).

We also present an estimate of heat production by friction between the ice sheet base and the subglacial stratum. This shear heating is the product of basal velocity and basal drag along flow. Yield strengths of till collected below the Whillans Ice Stream are a few kPa (Tulaczyk et al., 2000). Thus, the basal velocity approaches the surface velocity. In this calculation of the shear heat flux, we take basal velocity equal to the surface velocity, representing an upper bound on the shear heat flux. Since basal drag is poorly constrained, we calculate shear heat flux profiles using a range of basal drag values from 2 to 10 kPa.

Although we do not account for heat sources and sinks due to freezing or melting, and heat advection due to subglacial water flow, these are consistent with our calculated \(q_i\) which is an upper bound given \(T_b = T_m(p)\). This analytical approach neglects lateral ice advection, which may alter \(q_i\) within ice streams if lateral gradients in surface ice temperature are significant. However, along the Siple Coast surface temperature gradients are small (Comiso, 2000), and this analytical approach reproduces the ice temperature profile reasonably well at SLW (Fisher et al., 2015). A more thorough analysis of this source of variability would entail 3-D ice sheet modeling.

### 3. Results and Discussion

#### 3.1. GHF Observations

Two measurements of the thermal gradient at the WGZ show good agreement, yielding a temperature gradient of 0.050 \(\pm 0.004°C\) m\(^{-1}\) (mean \(\pm SD\)) (Figure 1; for the full record, see Figure S2 and Data Set S1).
The thermal conductivity ($k$) of sediments collected at the site range from 1.6 to 2.1 W m$^{-1}$°C$^{-1}$, with local variations that are likely associated with differences in grain size (Gangadhara Rao & Singh, 1999) (Figure S4), grain lithology, and/or porosity (Brigaud & Vasseur, 1989) (Figure 1c and Data Set S2). There is no clear trend in $k$ with depth, and we use the harmonic mean of measured $k$ values, 1.77 ± 0.15 W m$^{-1}$°C$^{-1}$, to calculate GHF.

At the WGZ, the vertical, conductive GHF is 88 ± 7 mW m$^{-2}$ (mean ± 1 SE, Table S1). The shallowest equilibrium sediment temperatures have the largest uncertainties (Figure 2b), perhaps because of disruption of shallow sediments by probe insertion. If these data are omitted, then the thermal gradient is ~18% greater, and GHF is 104 ± 3 mW/m$^2$. In contrast, the same tools and methods were applied at SLW, ~100 km away, yielding GHF of 285 ± 80 mW/m$^2$ (Fisher et al., 2015). An earlier measurement below the Ross Ice Shelf at J9, ~200 km from the WGZ, indicated GHF of 55 mW m$^{-2}$ (Figure 2) (Foster, 1978).

### 3.2. Processes Contributing to Elevated and Variable GHF in West Antarctica

There are a number of factors that can contribute to elevated and/or variable GHF, acting over a range of length scales (Table 1). We examine each of these factors to determine which could explain large variations in GHF (200 mW m$^{-2}$) over relatively short distances (<100 km), as observed below the Whillans Ice Stream. The spatial scales of crustal thickness variability are too broad and the magnitude of resulting GHF deviations too small to explain the observed GHF variability (Chaput et al., 2014; Fox Maule et al., 2005) (Text S1c and Figure 3a). Thermal conductivity variability can produce small-scale GHF variability by conductive refraction, but the maximum difference in GHF is 30 mW m$^{-2}$ (Text S1a). While variability in crustal radiogenic heat production can produce small-scale GHF variability as well, it is unlikely to enhance GHF by more than 18 mW m$^{-2}$ (Figure 2a and Text S1b) (Vilà et al., 2010). Erosion and lithospheric extension in West Antarctica produce small rates of vertical advection that enhance GHF by ≤10 mW m$^{-2}$ (Texts S1d and S1e) (Lachenbruch & Sass, 1978; Mancktelow & Grasemann, 1997).

Two remaining processes could generate the observed spatial variability in GHF: (a) recent magmatism at shallow crustal depths, and/or (b) advection of heat by crustal fluid flow, potentially associated with hydrothermal circulation. The influence of magmatic intrusions on GHF is estimated using the analytical, transient solution of Lachenbruch et al. (1976) for a prismatic intrusion (Figure 2b). In this model, the thermal conductivity of the surrounding crust is homogeneous and set to 2.8 W m$^{-1}$°C$^{-1}$, the initial temperature of the intrusion is set to 1,000°C, and the background GHF at the surface outside of the influence of the intrusion is set to 70 mW m$^{-2}$. GHF values in excess of 200 mWm$^{-2}$ are reached as a result of intrusions <5 km in diameter emplaced within the last 150 kyr. These intrusions can generate elevated GHF with spatial footprints less
Table 1
Observational Constraints on GHF Variability and Candidate Explanations

<table>
<thead>
<tr>
<th></th>
<th>Magnitude of GHF difference (mW m$^{-2}$)</th>
<th>Lateral extent of GHF difference (km)</th>
</tr>
</thead>
<tbody>
<tr>
<td>Observations (SLW-WGZ)</td>
<td>197 ± 85</td>
<td>108</td>
</tr>
<tr>
<td>Observations (WGZ-J9)</td>
<td>33 ± 7</td>
<td>228</td>
</tr>
<tr>
<td>Candidate explanations</td>
<td></td>
<td></td>
</tr>
<tr>
<td>Hydrothermal circulation</td>
<td>1000s</td>
<td>0.1–100s</td>
</tr>
<tr>
<td>Magmatic intrusion</td>
<td>1000s</td>
<td>&lt;10</td>
</tr>
<tr>
<td>Crustal thickness variability</td>
<td>≤60</td>
<td>&gt;130</td>
</tr>
<tr>
<td>Thermal conductivity</td>
<td>&lt;30</td>
<td>&gt;1</td>
</tr>
<tr>
<td>Radiogenic heat production</td>
<td>≤18</td>
<td>&lt;20</td>
</tr>
<tr>
<td>Lithospheric extension</td>
<td>≤10</td>
<td>≥75</td>
</tr>
<tr>
<td>Erosion</td>
<td>&lt;4</td>
<td>10–200</td>
</tr>
</tbody>
</table>

Figure 2. (a) GHF measurements and estimates for West Antarctica (Engelhardt, 2004a; Fisher et al., 2015; Foster, 1978; Fudge et al., 2013) and the western Ross Sea region (Morin et al., 2010, and references therein; Schröder et al., 2011) overlain on ice velocity (Rignot et al., 2011). Grounding line outlined black (Bindschadler et al., 2015). Profile line (A–A') shown in black. Extent of GHF estimates below Thwaites glacier (THW, dashed line) (Schroeder et al., 2014). (b) Estimates of spatial variability in heat conduction and production along the profile line shown in Figure 2a, as difference from mean conductive heat flux along that profile (79 mW m$^{-2}$). (c) Shear heat flux estimates calculated from ice velocity and associated errors. GHF measurements and estimates close to the profile line are plotted (mean ±1 SE, SLW value lies off axis).
Dotted lines show the envelope of ±15% variation in crustal thermal conductivity as a function of magnetic crustal thickness. The SLW value lies well above the plot and gray lines) compared with GHF measurements and estimates (blue) as a function of GHF models for West Antarctica (An et al., 2015; FM=Fox et al., 2005; SR=Shapiro & Ritzwoller, 2004) and GHF measurements in the Basin and Range Province (Berg et al., 1989) (Figure S9), a range that overlaps with the geothermal gradient of 91–162°C km⁻¹ measured at SLW (Fisher et al., 2015).

The flow of crustal fluids can also increase GHF within a broad area or redistribute heat locally, depending on fluid pathways, flow rates, and the depth of circulation (Fisher & Harris, 2010). Hydrothermal circulation in basement rocks, even below sediments, can generate GHF anomalies with spatial scales of several to tens of kilometers (e.g., Davis et al., 1997; Fisher et al., 1990). Vigorous local convection can lead to isothermal conditions in a buried aquifer, resulting in large differences in GHF (several hundred mW m⁻²) through overlying strata as a function of depth to the aquifer top (Davis et al., 1997; Spinelli & Fisher, 2004). Where basement is exposed at the base of the ice, it may provide a conduit for discharge and recharge of hydrothermal fluids, increasing and decreasing GHF, respectively (e.g., Davis et al., 1992; Fisher et al., 2003; Villinger et al., 2002). The magnitude of GHF anomalies where basement outcrops at the surface can be several W m⁻², relative to background values of ~100 mW m⁻² (e.g., Davis et al., 1992; Fisher et al., 2003; Villinger et al., 2002). The gravity data collected at WGZ suggests that basement topography may exist (Muto et al., 2013), but there is no such evidence at SLW. The gravity data are consistent with a crustal fault, which could enhance permeability by several orders of magnitude relative to unfaulted bedrock (Seront et al., 1998), focusing vertical fluid advection and elevating GHF. Thus, either magmatism or advection of heat by fluids may contribute to high and spatially variable GHF in West Antarctica.

These two processes have also played a role in generating GHF variability in other rift systems (Reiter et al., 1975). The observed variability of the GHF in West Antarctica is consistent with that of other rift systems (Davies & Davies, 2010) such as the Basin and Range Province of North America, which is often considered to be a geologic analog for the West Antarctic Rift System (An et al., 2015; Behrendt et al., 1994; Decesari et al., 2007) suggests that rifting-related processes such as magmatism or preferential advection of crustal fluids may affect ice dynamics by enhancing GHF.

3.3. Implications of High and Variable GHF for Slow-Flowing Ice

Given that GHF measurements reveal a wide range of variability, from tens of mW m⁻² over distances of ~200 km (WGZ, J9, SD) to ~200 mW m⁻² over ~100 km (WGZ, SLW), we compare this variability with
independent estimates for the variability in heat flux on the Siple Coast (Figure 2b).

Estimated lateral variations in the vertical conductive heat flux are dominated by spatial variations in ice thickness. Calculated fluxes increase by 7 mW m⁻² per 100 m decrease in ice thickness, resulting in spatial variations of 7–28 mW m⁻² over 100 km from interstream ridge to ice stream trough. In contrast, estimated lateral variations in the vertical conductive heat flux due to changes in accumulation rate are generally <10 mW m⁻² per 100 km. Estimated lateral variations in the vertical conductive heat flux due to changes in surface temperature are generally <5 mW m⁻² per 100 km. The frictional heat flux due to ice sliding over subglacial sediments is poorly constrained due to uncertainties in basal resistance and basal sliding velocity, but is estimated to be <125 mW m⁻² near WGZ where ice velocity is around 300 m yr⁻¹. These sources of variability in heat flux are less than the spatial variability in GHF of ~200 mW m⁻² per 100 km (WGZ, SLW) and of the same magnitude as the spatial variability between other GHF estimates (WGZ, J9, SD).

4. Conclusions

Current geophysical GHF models underestimate the observed magnitude and spatial variability of GHF, which may be enhanced by magmatism or advection of crustal fluids. Large differences in sea level rise predictions from Antarctica result from two GHF models with narrow GHF distributions (Bougamont et al., 2015). The observed spatial variability in GHF raises the possibility that GHF plays a greater role in ice dynamics than generally considered. Zones of elevated GHF below the WAIS can produce considerable volumes of subglacial meltwater (Vogel & Tulaczyk, 2006) and may contribute to the development and dynamics of subglacial lakes, the advection of organic and inorganic compounds into subglacial habitats, and thus the presence and metabolism of microbial biomes (Christner et al., 2014; Jørgensen & Boetius, 2007). Seroussi et al. (2017) found that locally high GHF (≥150 mW m⁻²) below the Whillans Ice Stream was needed to reproduce the observed subglacial lakes in an ice sheet model. As the ice sheet thins, increasing the vertical conductive heat flux, GHF variability may be more important to predictions of the basal thermal regime, particularly the development of basal frozen zones such as ice rises that might stabilize ice retreat (Favier & Pattyn, 2015; Rignot et al., 2004).

Bed topography and ice sheet thickness are relatively well constrained for much of West Antarctica (Fretwell et al., 2013). Spatial variability in GHF may contribute more to the uncertainty in the basal thermal regime of West Antarctica than does the remaining uncertainty in ice thickness, which is equivalent to GHF uncertainty of 4 mW m⁻² along the Siple Coast (Table S6). More direct GHF observations are needed to constrain continental GHF models. Ice sheet modeling could direct GHF observations to locations where future ice sheet mass balance is most sensitive to GHF, to maximize the impact of field measurements. Until such observational constraints become available, we recommend running ensembles of ice sheet models for multiple spatial distributions of GHF below the WAIS, including distributions as broad as that in the Basin and Range Province, to set more realistic limits on rates of ice loss.

References


Acknowledgments

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